

SIM Metrology School, NIST, Gaithersburg, Oct. 28-Nov.1, 2013

Optical Metrology – Photometry

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Physical Measurement Laboratory

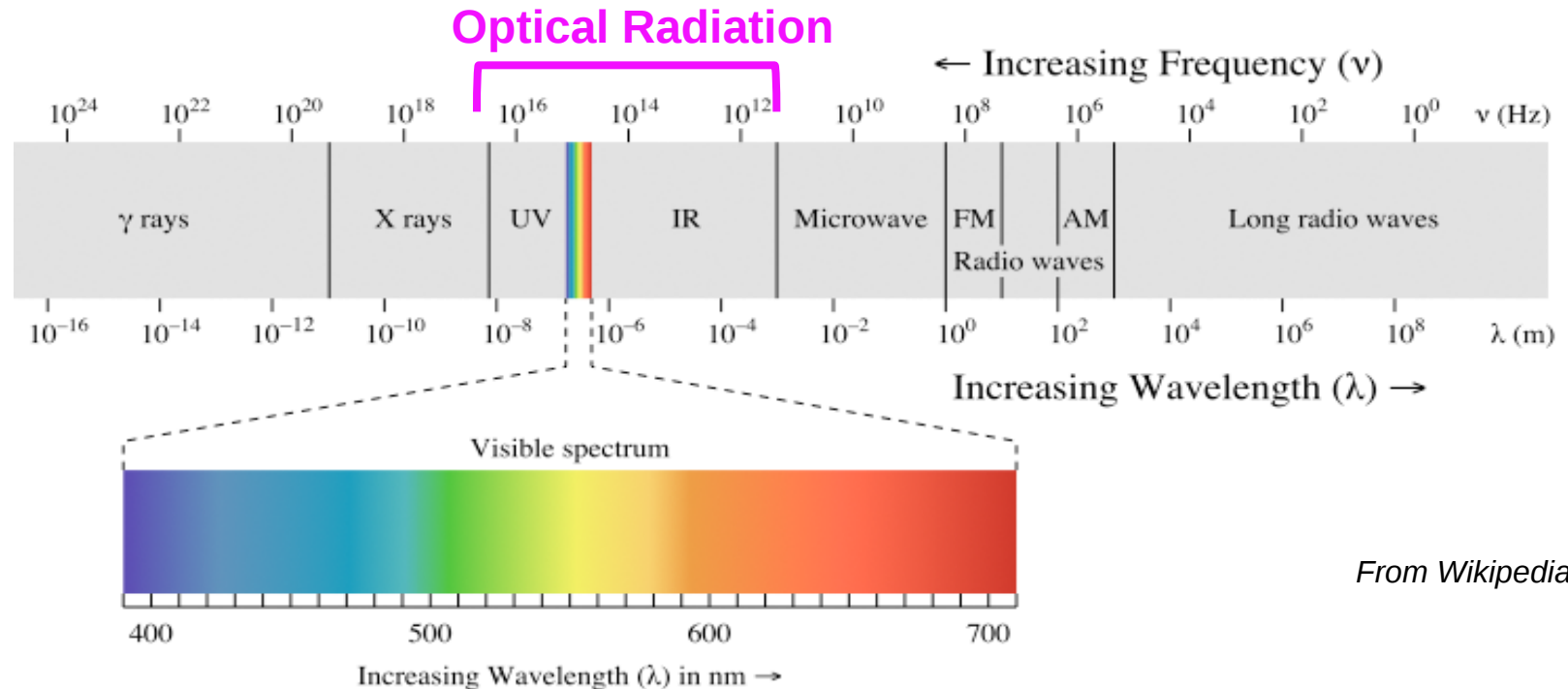
National Institute of Standards & Technology

OUTLINE

1. Optical Radiation
2. Radiometry and photometry
3. Quantities and units
4. Human visual response
5. Colorimetry
6. Standards organizations
7. Laboratory Examples

Optical Radiation

Electromagnetic Radiation

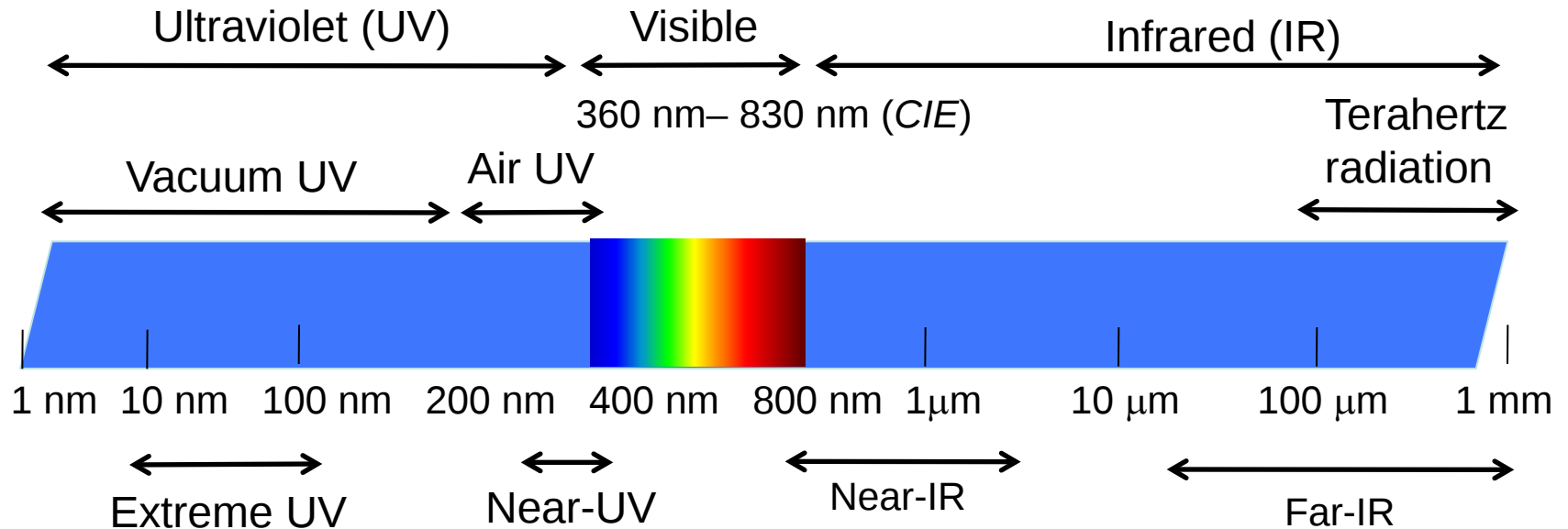


From Wikipedia

Optical Radiation

Electromagnetic radiation at wavelengths between the region of transition to X-rays ($\lambda \approx 1$ nm) and the region of transition to radio waves ($\lambda \approx 1$ mm) *International Lighting Vocabulary (CIE 17.4)*

Optical Radiation



UV-A: 315 nm to 400 nm

UV-B: 280 nm to 315 nm

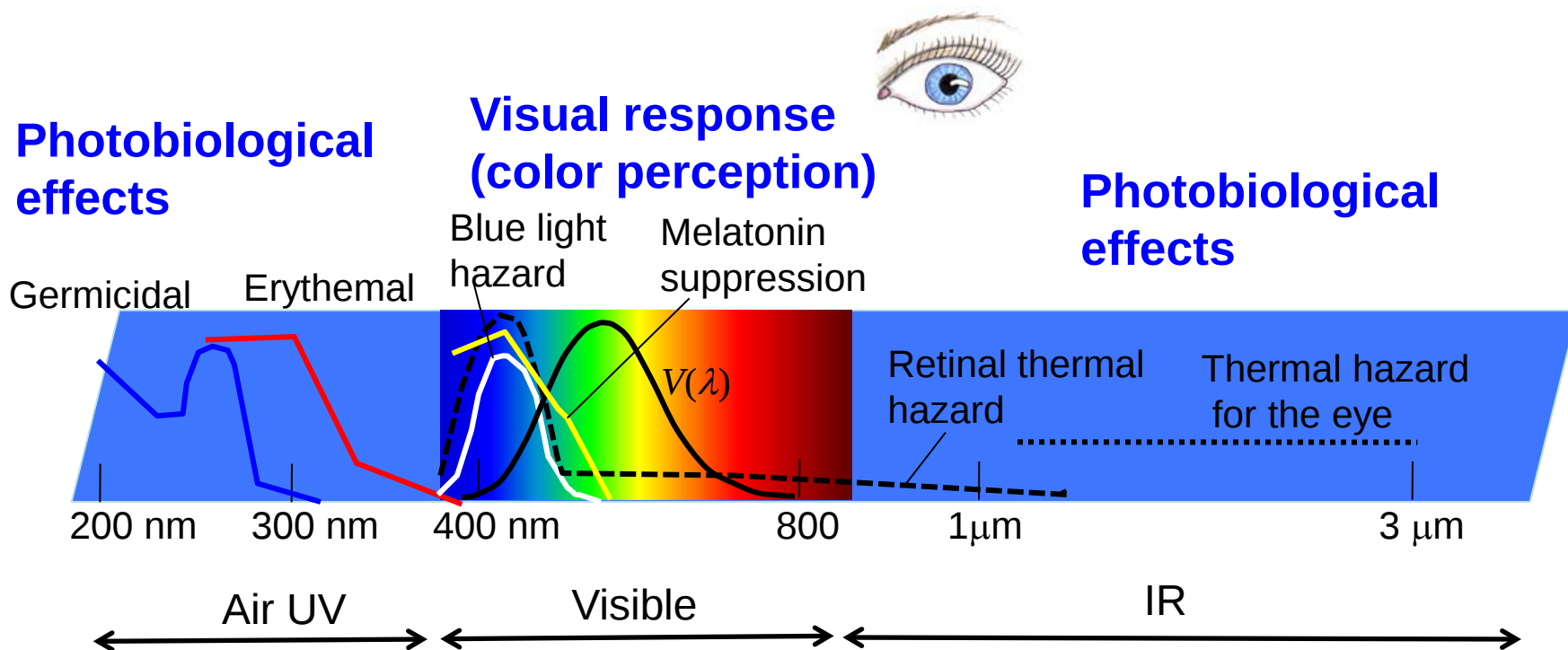
UV-C: 100 nm to 280 nm (CIE)

IR-A: 780 nm to 1400 nm

IR-B: 1,4 μm to 3,0 μm;

IR-C: 3 μm to 1 mm. (CIE)

Physiological Effects of Optical Radiation



Action spectrum: Effectiveness of photochemical or biological reaction of optical radiation as a function of wavelength.

Measurements of Optical Radiation

Radiometry

(all wavelengths, purely physical)

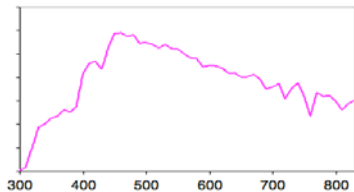


Radiant power
(radiant flux)

[watt] or [W]

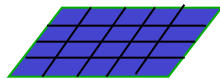
Radiant energy

[joule] or [J]



Spectral irradiance

[W m⁻² nm⁻¹]



Irradiance

[W/m²]

Radiance

[W sr⁻¹ m⁻²]



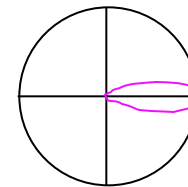
Luminous flux

[lumen] or [lm]



Illuminance

[lux] or [lx] (= lm/m²)



Luminous intensity

[candela] or [cd]

(= lm/sr)



Luminance

[cd/m²]

The Candela

Unit of luminous intensity

SI-base units

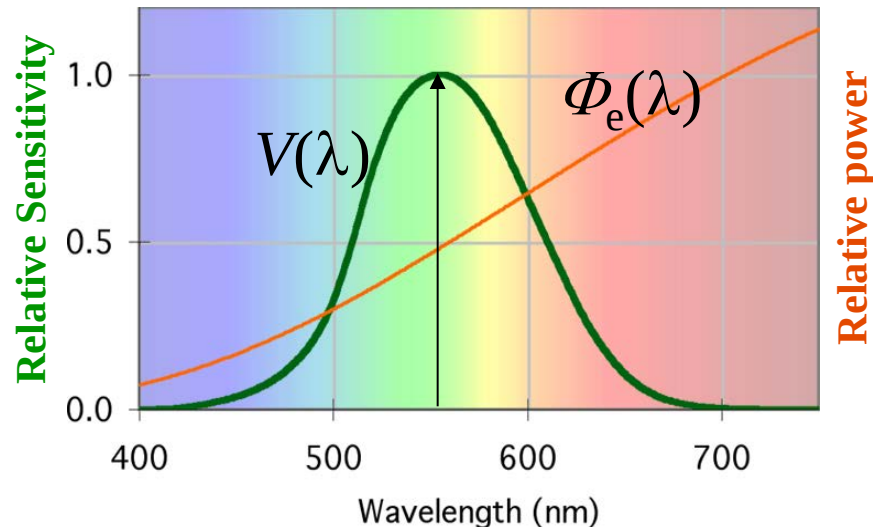
meter
second
kilogram
ampere
candela
kelvin
mole

“The candela is the luminous intensity, in a given direction, of a source that emits monochromatic radiation of frequency 540×10^{12} hertz and that has a radiant intensity in that direction of $(1/683)$ watt per steradian.” (CGPM 1979)

$$1 \text{ cd} = 1/683 \text{ [W/sr]} \quad \text{at } 555.016 \text{ nm}$$
$$(1 \text{ W} = 683 \text{ lm}) \quad (\text{in std. air})$$



Luminous efficiency $V(\lambda)$ (CIE 1924)



Photometric quantity

$$\Phi_v = K_m \int_{\lambda} \Phi_{e,\lambda} V(\lambda) d\lambda$$

$$K_m \approx 683 \text{ [lm/W]}$$

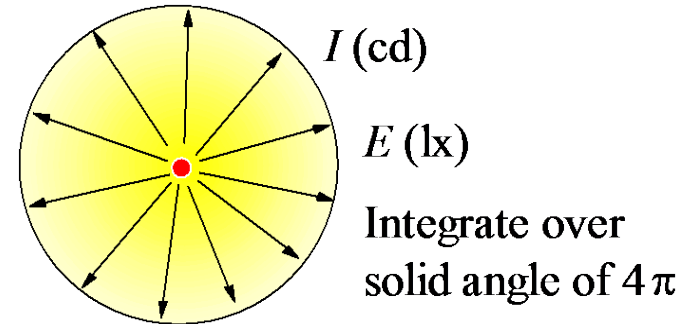
(BIPM 1983)

The Lumen

Unit of (total) luminous flux

Total Luminous Flux [lm]

The geometrically-total luminous flux of a light source. It is given by the spatial integration of luminous intensity I [cd]



$$\Phi = \int_{\Omega} I \, d\Omega$$

or the integration of illuminance E [lx]:

$$\Phi = \int_A E \, dA$$

(A : area of a closed surface on which illuminance is given)

Indoor/Outdoor INTERIOR/EXTERIOR	
Light output Rendimiento luminoso	1600 Lumens Lúmenes
Energy used Energía usada	23 Watts Vatios
Life Vida útil	8000 Hours Horas



$$1600 \text{ lm} / 23\text{W} = 70 \text{ lm/W}$$

Why photometry so important? (example)

LED lighting
products
Market



Quality assurance

Certify products

Regulations

Non-standard LED lamps

min. 50 lm/W, 200 lm/lamp

Replacement lamps

a) omnidirectional lamps

min. 50 lm/W, life 25,000 h

b) Decorative lamps

min. 40 lm/W)

Color rendering requirements

CRI $R_a \geq 80$, $R_9 \geq 0$

CCT 2700, 3000, 3500, 4000 K

Nominal wattage of lamp to be replaced (watts)	Minimum initial light output of LED lamp (lumens)
25	200
35	325
40	450
60	800
75	1,100
100	1,600



Product
performance
data measured



Accredited
Manufacturers,
testing labs

Testing
Accreditation
Program

Proficiency Testing

Test Method Standard
(IES LM-79 in USA)

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Quantities and Units for Optical Radiation

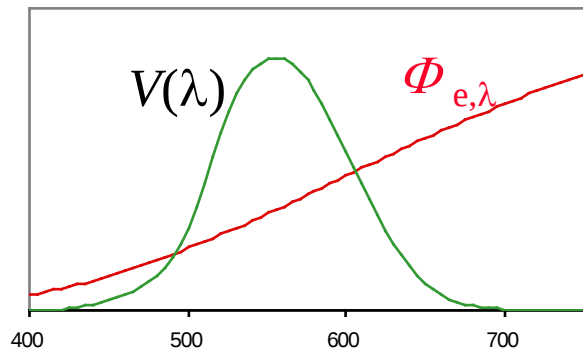
Radiometric quantity	Unit	Photometric quantity	Unit	relationship with lumen
Radiant flux	W (watt)	Luminous flux	lm (lumen)	
Radiant intensity	W/sr	Luminous intensity	cd (candela)	$\text{lm} \cdot \text{sr}^{-1}$
Irradiance	W/m^2	Illuminance	lx (lux)	$\text{lm} \cdot \text{m}^{-2}$
Radiance	$\text{W} \cdot \text{sr}^{-1} \cdot \text{m}^{-2}$	Luminance	cd/m^2	$\text{lm} \cdot \text{sr}^{-1} \cdot \text{m}^{-2}$
Radiant exitance	W/m^2	Luminous exitance	lm/m^2	$\text{lm} \cdot \text{m}^{-2}$
Radiant energy	J (joule)	Luminous energy	$\text{lm} \cdot \text{s}$	
Radiant exposure	J/m^2	Luminous exposure	$\text{lx} \cdot \text{s}$	$\text{lm} \cdot \text{s} \cdot \text{m}^{-2}$
Total radiant flux	W	Total luminous flux	lm	
Radiance temperature	K (kelvin)	Color temperature	K (kelvin)	

Luminous Flux

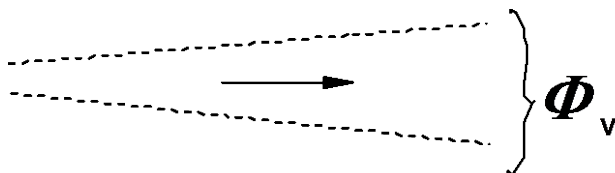
Unit: **lumen**

Symbol for unit: **lm**

Symbol for quantity: Φ , Φ_v



$$\Phi_v = K_m \int_{\lambda} \Phi_{e,\lambda} V(\lambda) d\lambda$$



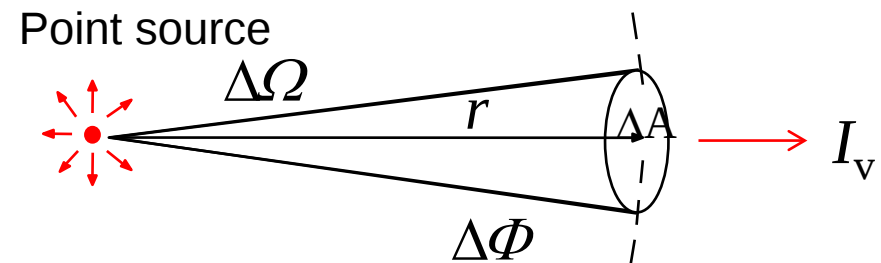
Luminous Intensity

(of a source, in a given direction)

Unit: **candela** = lm/sr

Symbol for unit: **cd**

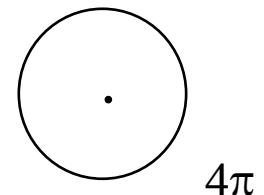
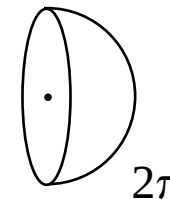
Symbol for quantity: I , I_v



$$I_v = \frac{d\Phi_v}{d\Omega}$$

Solid angle

$$\Delta\Omega = \Delta A / r^2 \text{ [sr]}$$

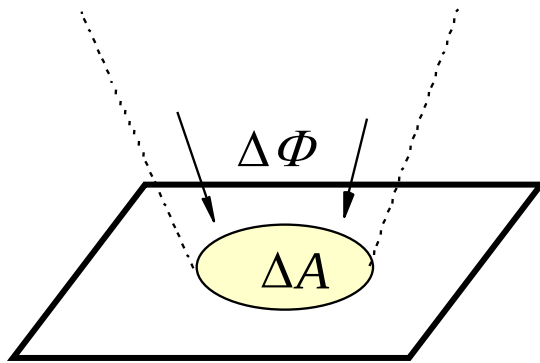


Illuminance (at a point of a surface)

Unit: **lux** = lm/m²

Symbol for unit: **lx**

Symbol for quantity: **E , E_v**



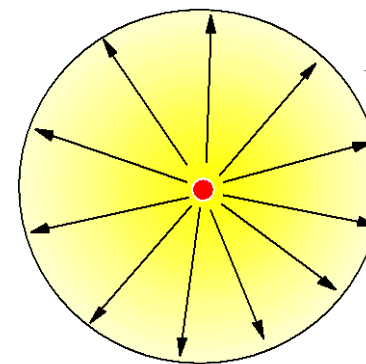
$$E_v = \frac{d\Phi_v}{dA}$$

Total Luminous Flux (of a source)

Unit: **lumen**

Symbol for unit: **lm**

Symbol for quantity: **Φ , Φ_v**



I (cd)

E (lx)

Integrate over
solid angle of 4π

$$\Phi = \int_{\Omega} I \, d\Omega$$

or

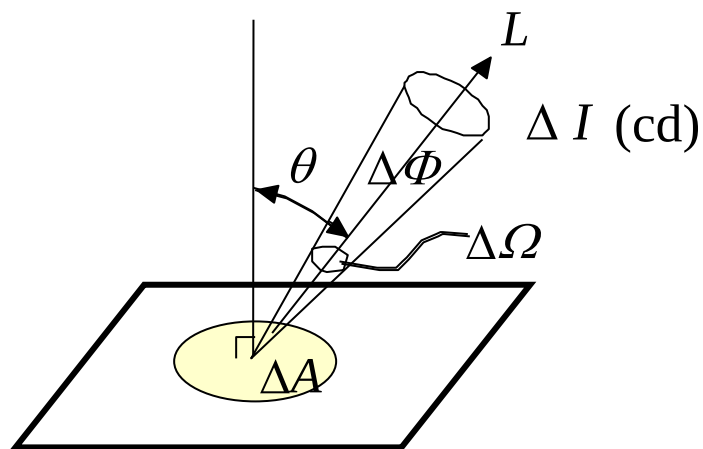
$$\Phi = \int_A E \, dA$$

(A: area of a closed
surface on which
illuminance is given)

Luminance (at a point of a surface, in a given direction)

Unit: **cd/m²** (= lm sr⁻¹ m⁻²)

Symbol for quantity: ***L***, ***L_v***



$$L_v = \frac{d\Phi_v}{d\Omega dA \cos \theta}$$

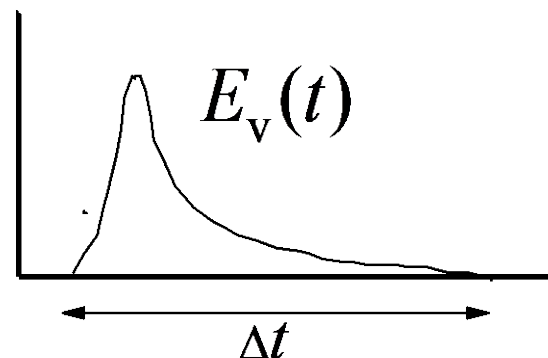
$$L_v = \frac{\Delta I}{\Delta A \cos \theta}$$



Luminous exposure

Unit: **lux · second** (lx · s)

Symbol for quantity: ***H***, ***H_v***



$$H_v = \int_{\Delta t} E_v(t) dt$$

Luminous energy

Unit: **lumen · second** (lm · s)

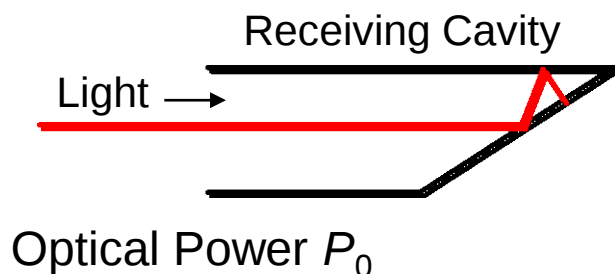
Symbol for quantity: ***Q***, ***Q_v***

$$Q_v = \int_{\Delta t} \Phi_v(t) dt$$

Primary Standards for Optical Radiation

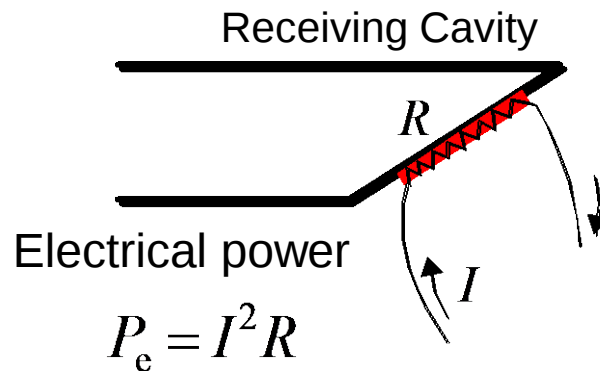
(1) Electrical Substitution Radiometer

Optical Heating



P_0 creates temperature increase ΔT in the cavity

Electrical Heating



P_e creates the same ΔT in the cavity

The optical power is measured as equal to the electrical power.

Uncertainty of room-temperature ESR is $\sim 0.1\%$ ($k=2$)

NIST Cryogenic Radiometer

Primary Optical Watt Radiometer (POWR)

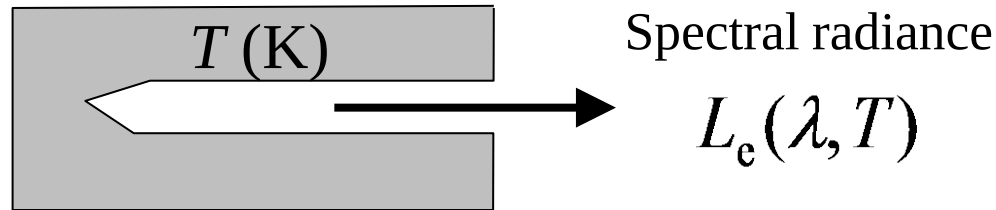


- Liquid helium/nitrogen cryostat
- Cavity cooled to 4 K
- Uncertainty $\sim 0.01\%$ ($k=2$); 200 nm to ~ 2500 nm
- Basis of all NIST radiometric and photometric units.



Detector module

(2) Blackbody radiation



Planck's Law

$$L_e(\lambda, T) = c_1 n^{-2} \pi^{-1} \lambda^{-5} \left[\exp\left(\frac{c_2}{n\lambda T}\right) - 1 \right]^{-1}$$

$$c_1 = 2\pi h c^2 = 3.74177153 \times 10^{-16} \text{ W} \cdot \text{m}^2$$

$$c_2 = hc/k = 1.438770 \times 10^{-2} \text{ m} \cdot \text{K}$$

(2010 CODATA)

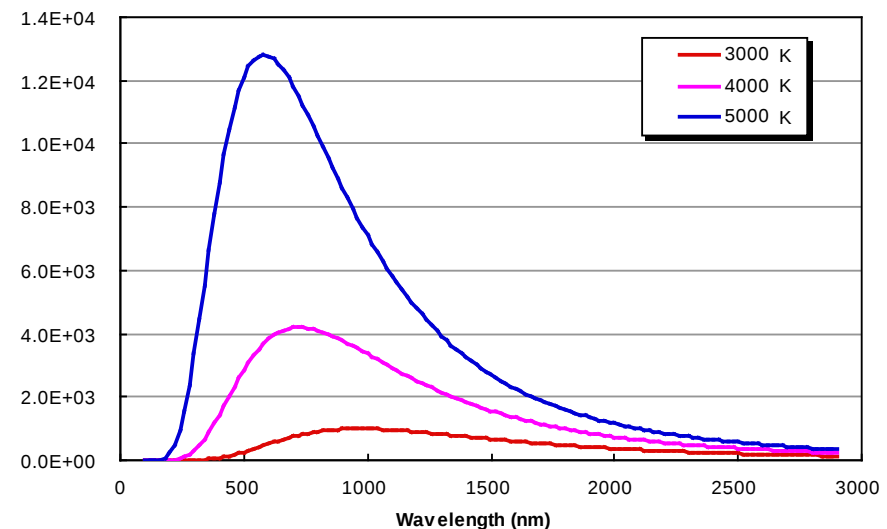
h : Planck constant

c : speed of light in vacuum

k : Boltzmann constant

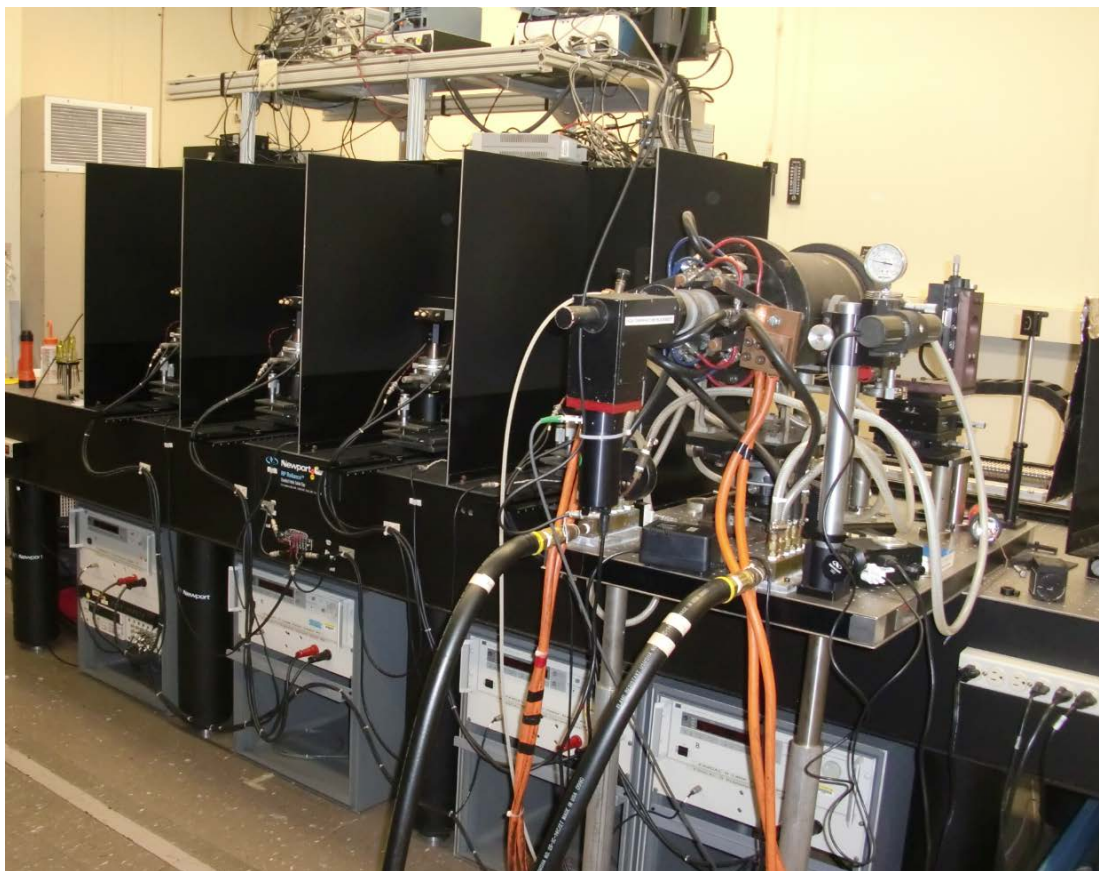
n ($=1.00028$): the refractive index of standard air

λ : wavelength



NIST Blackbody Radiation Facility

FASCAL 2 (Facility for Automated Spectral Irradiance Calibrations)



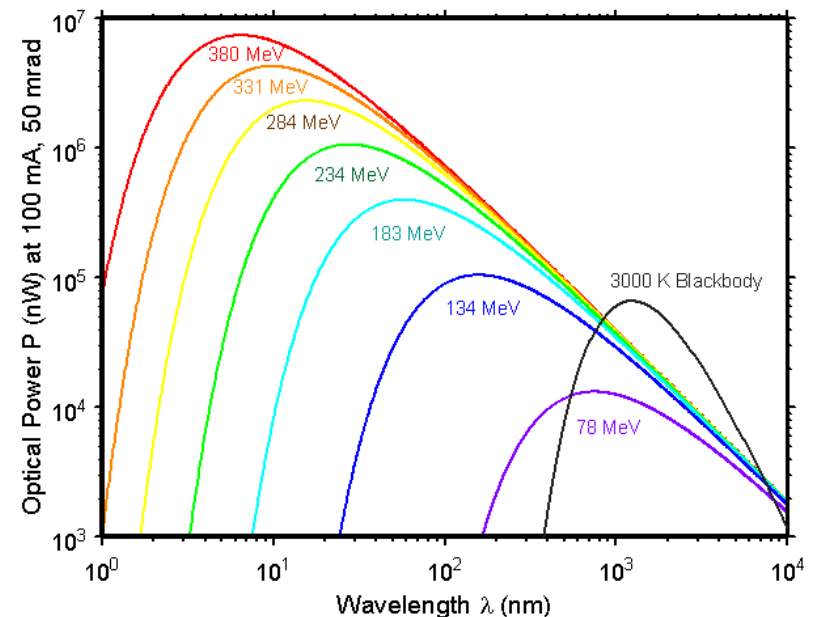
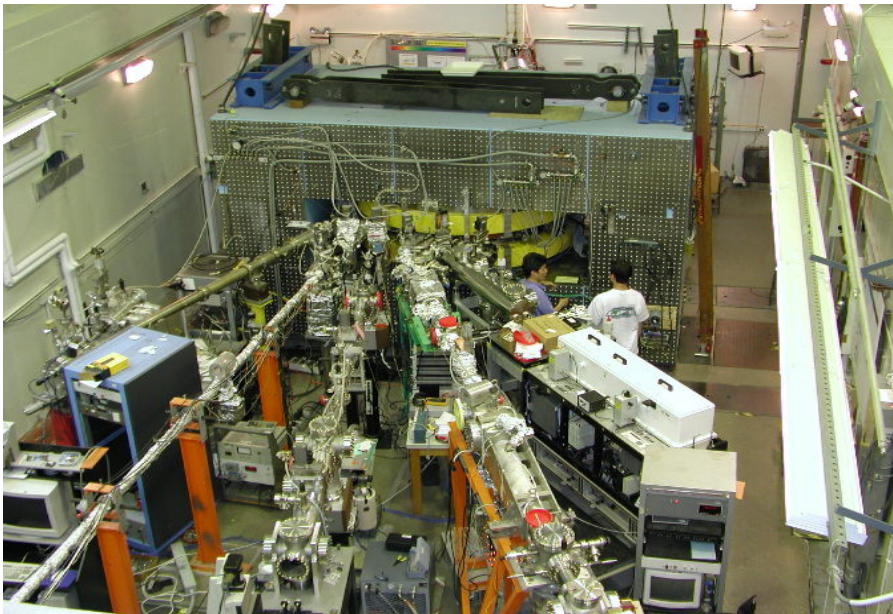
- Basis for NIST spectral irradiance scale 250 nm to 2400 nm.
- High temperature blackbody 2950 K.
- 1000 W FEL lamps issued.
- The scale provided to NMIs worldwide.

(3) Synchrotron radiation

Spectral radiant flux from a beam line is given by Schwinger equation.
Synchrotron radiation is used primarily for UV metrology.

NIST SURF III

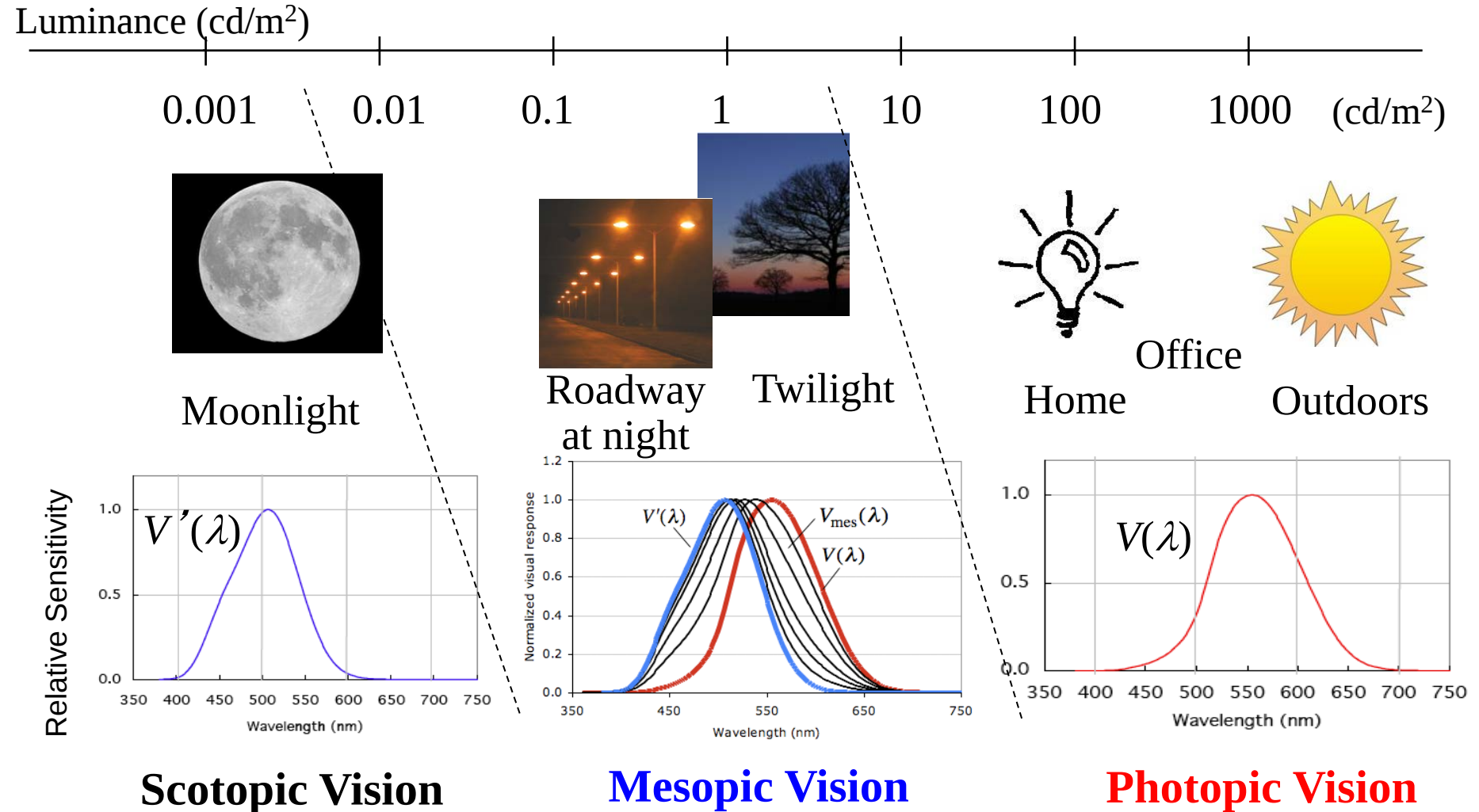
Synchrotron Ultraviolet Radiation Facility



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Human Visual Response



Photometric Quantities

CIE 18.2 The Basis of Physical Photometry (1983)

BIPM Monographie – Principles Governing Photometry (1983)

Luminous flux

For photopic vision:

$$\Phi_v = K_m \int_{\lambda} \Phi_{e,\lambda} V(\lambda) d\lambda$$

$$K_m = 683 \text{ lm/W (approx.)}$$

For scotopic vision:

$$\Phi_v = K'_m \int_{\lambda} \Phi_{e,\lambda} V'(\lambda) d\lambda$$

$$K'_m = 1700 \text{ lm/W (approx.)}$$

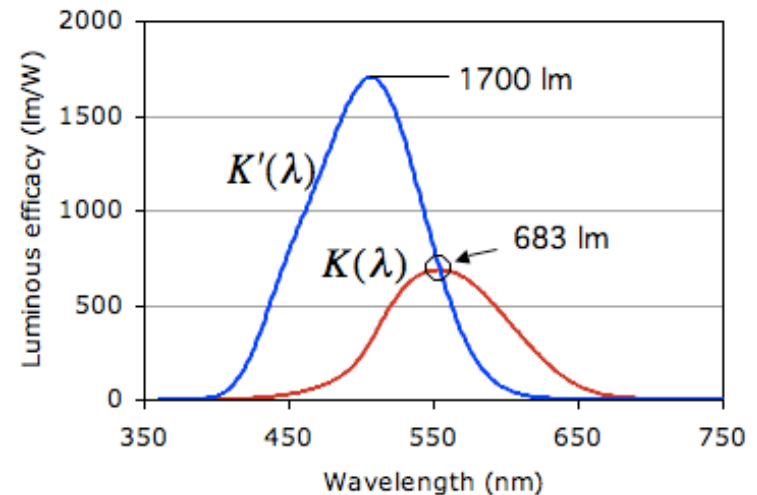
(Scotopic vision is rarely used.)

Spectral luminous efficacy

$$K(\lambda) = K_m V(\lambda) \quad \text{for photopic}$$

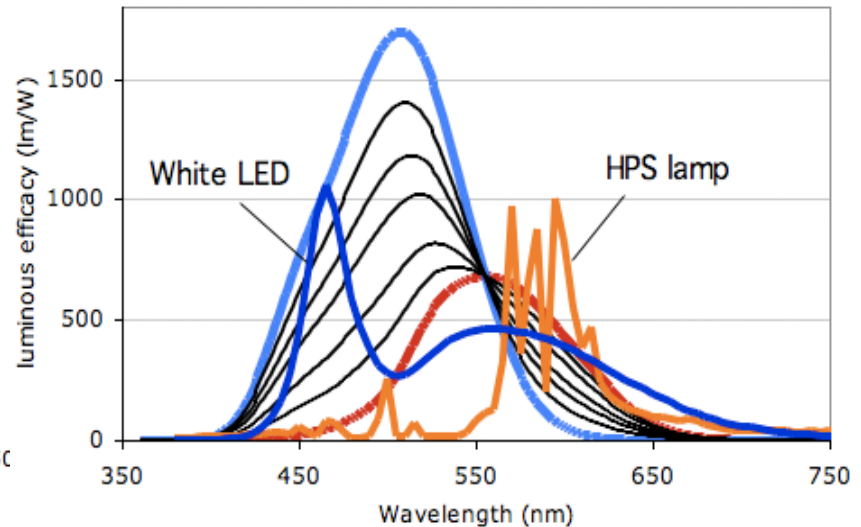
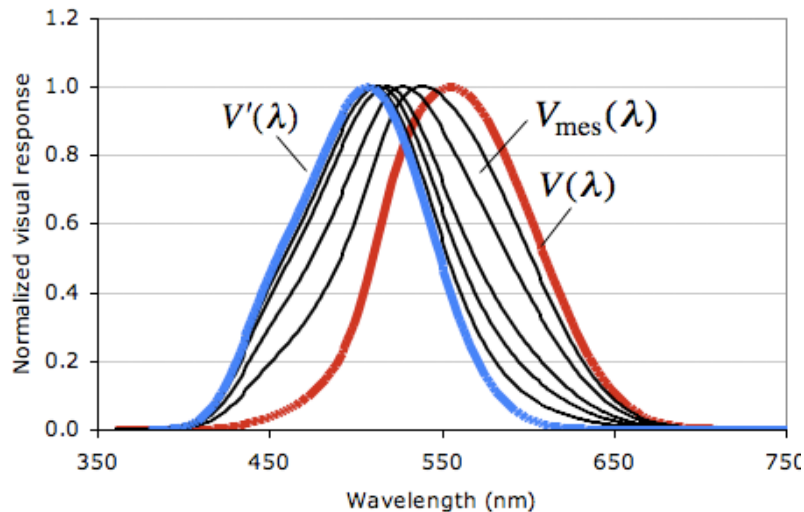
$$K'(\lambda) = K'_m V'(\lambda) \quad \text{for scotopic}$$

$K(\lambda) = K'(\lambda) = 683 \text{ lm/W}$ at the frequency $540 \times 10^{12} \text{ Hz}$



Mesopic Photometry

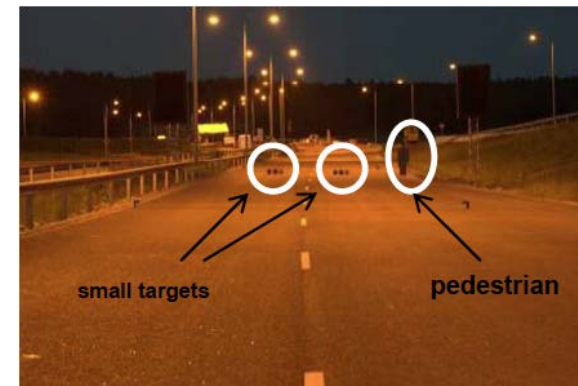
CIE 191:2010: Recommended System for Mesopic Photometry based on Visual Performance



$$M(m)V_{mes}(\lambda) = m V(\lambda) + (1-m)V'(\lambda) \quad \text{for } 0 \leq m \leq 1$$

$$\text{and } L_{mes} = \frac{683}{V_{mes}(\lambda_0)} \int V_{mes}(\lambda) L_e(\lambda) d\lambda$$

m is determined by adaptation luminance of the field.



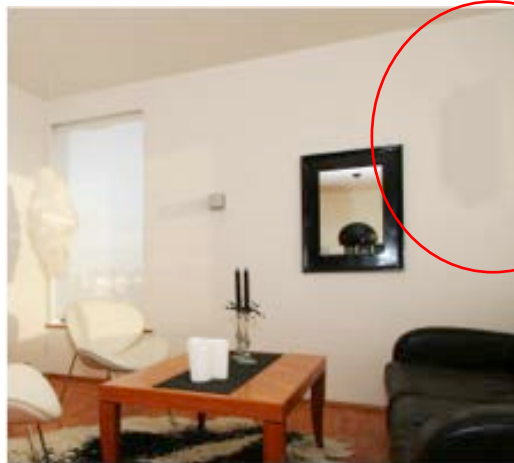
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Why accurate color so important?



Colors of parts do not match

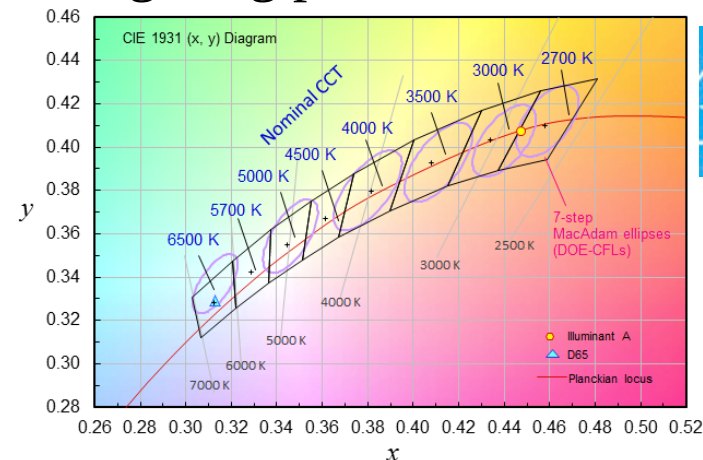


Color of paint does not match though same spec.



Skin tone does not appear right.

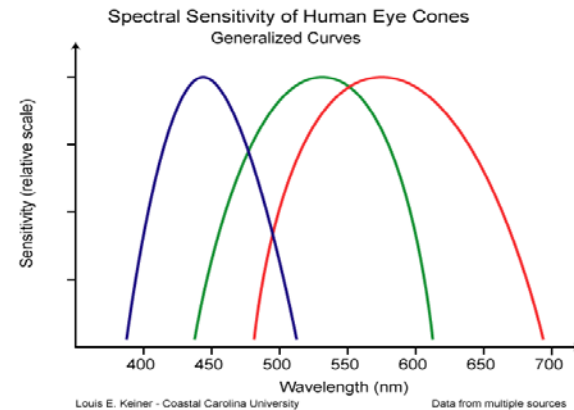
ANSI standard on LED lighting products



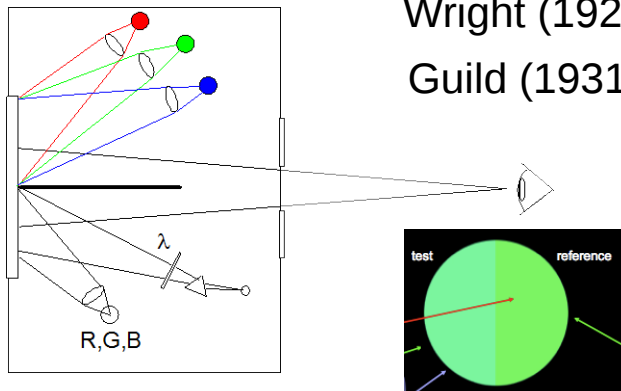
Color Vision



Trichromacy theory (Young, 1800' s)
3 cone photopigments in retina

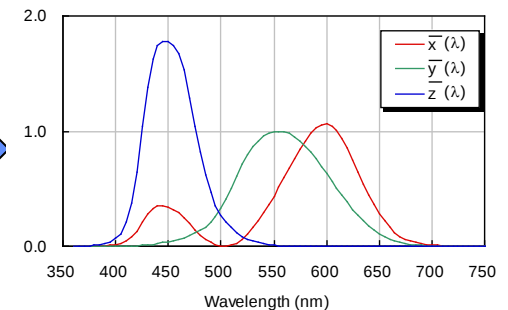
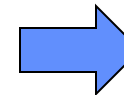
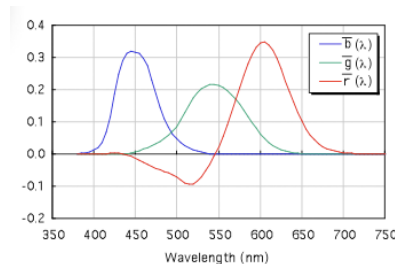


Color matching experiments



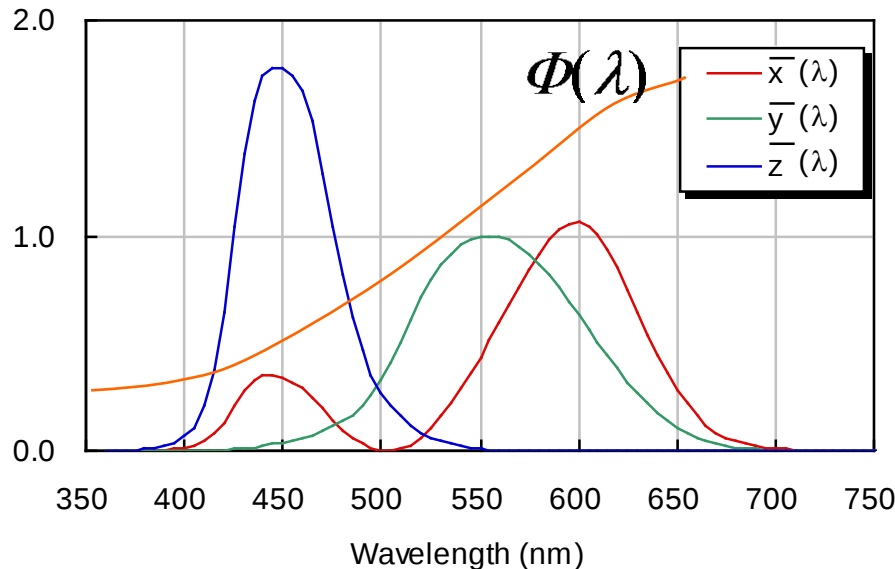
CIE 1931 Color Matching Functions

Results



1931 CIE System of Colorimetry

CIE 1931 Color Matching Functions



Tristimulus Values

$$X = k \int_{\lambda} \Phi(\lambda) \bar{x}(\lambda) d\lambda$$

$$Y = k \int_{\lambda} \Phi(\lambda) \bar{y}(\lambda) d\lambda$$

$$Z = k \int_{\lambda} \Phi(\lambda) \bar{z}(\lambda) d\lambda$$

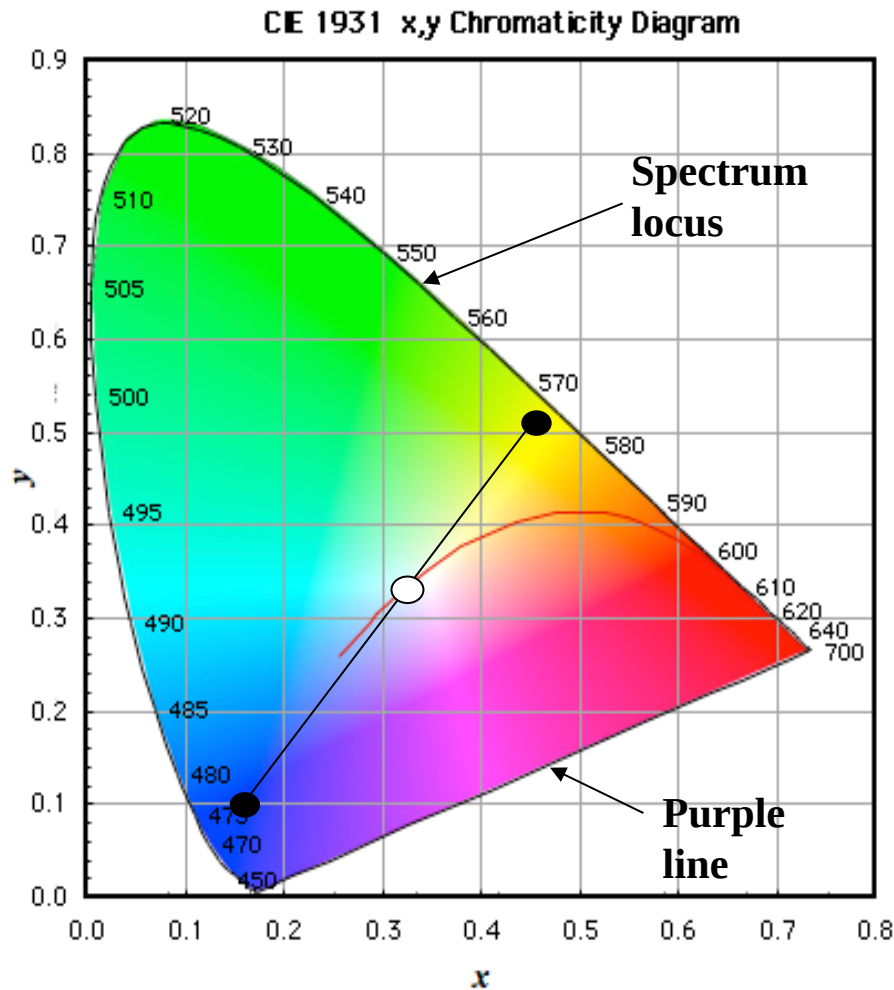
$$\bar{y}(\lambda) \equiv V(\lambda)$$

CIE 15:2004 Colorimetry, 3rd ed.

ISO 11664-1:2008(E)/CIE S 014-1/E:2006 Standard Colorimetric Observer

BIPM Monographie – Principles Governing Photometry (1983)

Chromaticity Coordinates

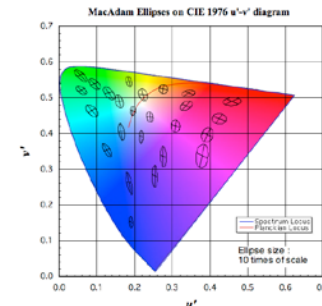


Chromaticity Coordinate
(x , y)

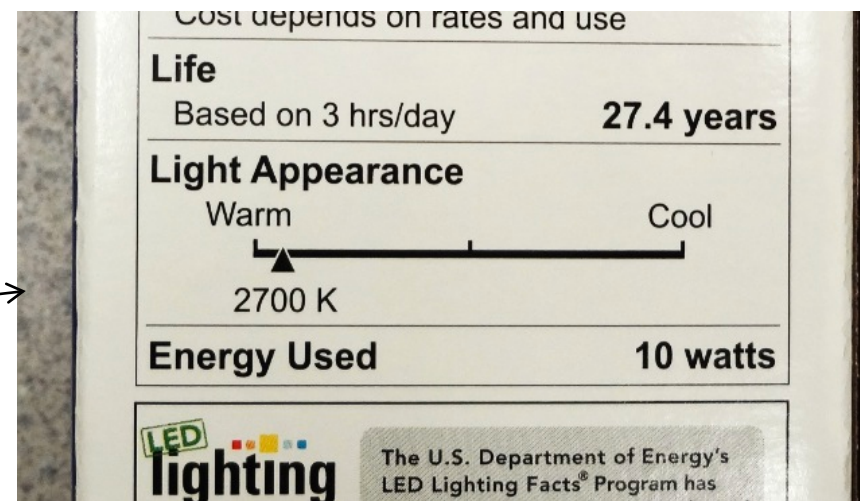
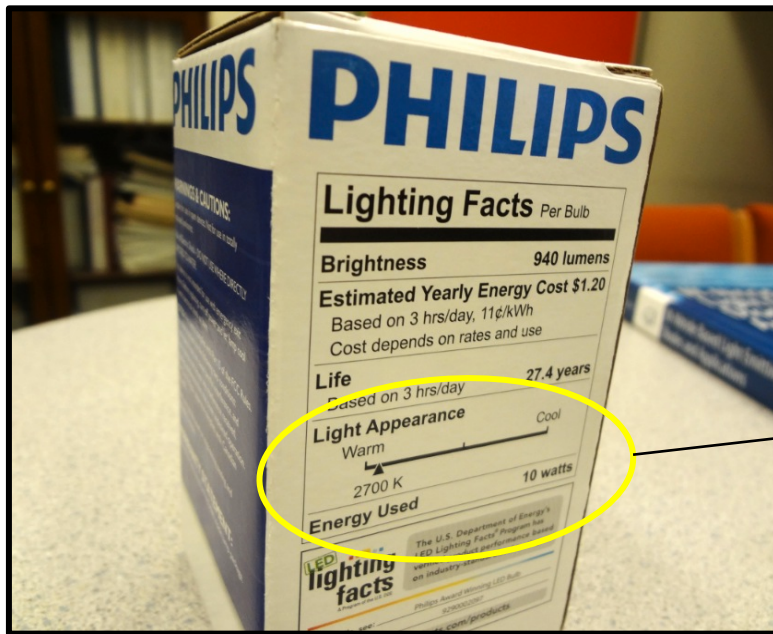
$$x = \frac{X}{X + Y + Z}$$

$$y = \frac{Y}{X + Y + Z}$$

CIE 1931 (x , y) diagram
CIE 1976 (u' , v') diagram

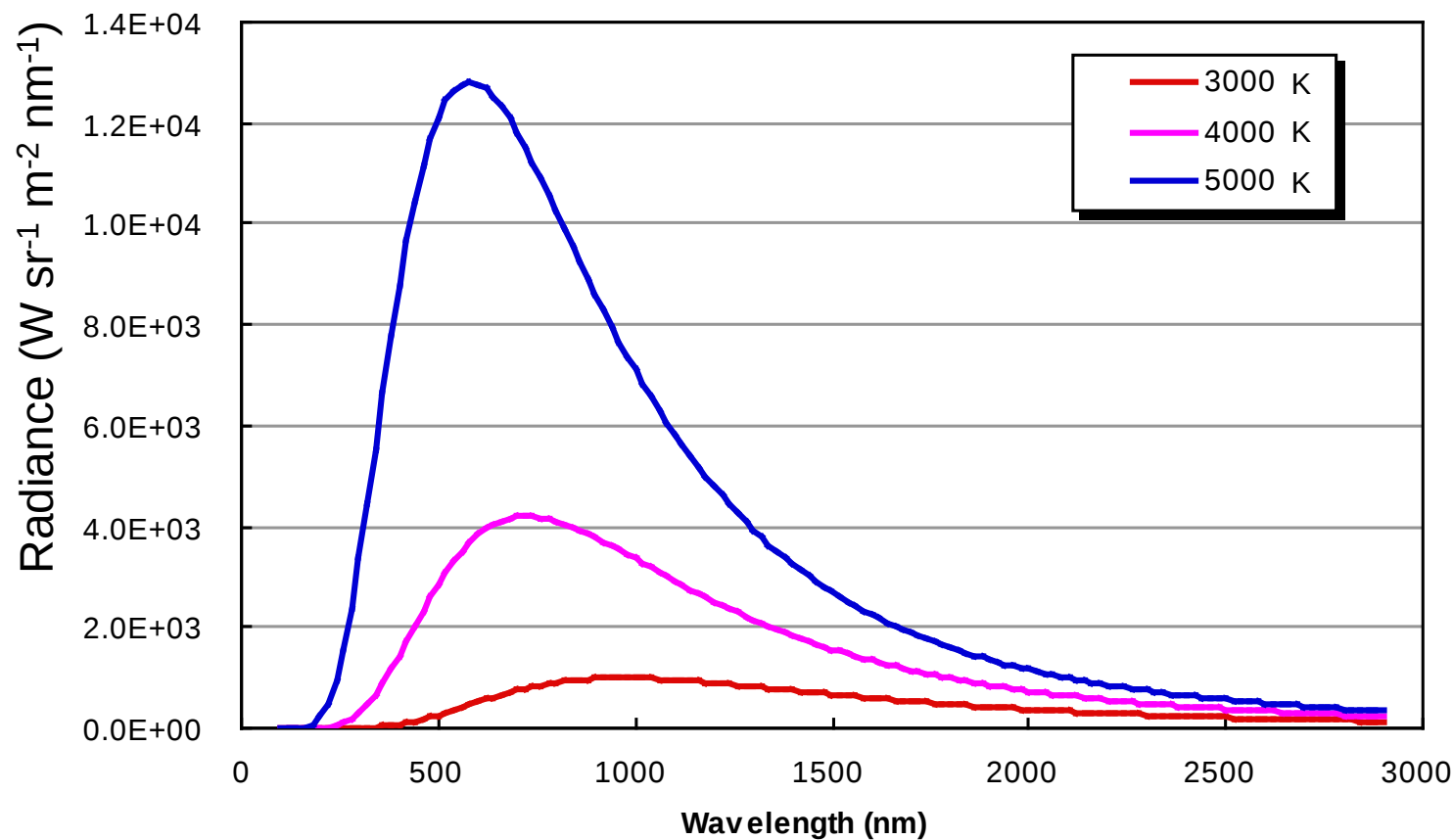


Color specification in products

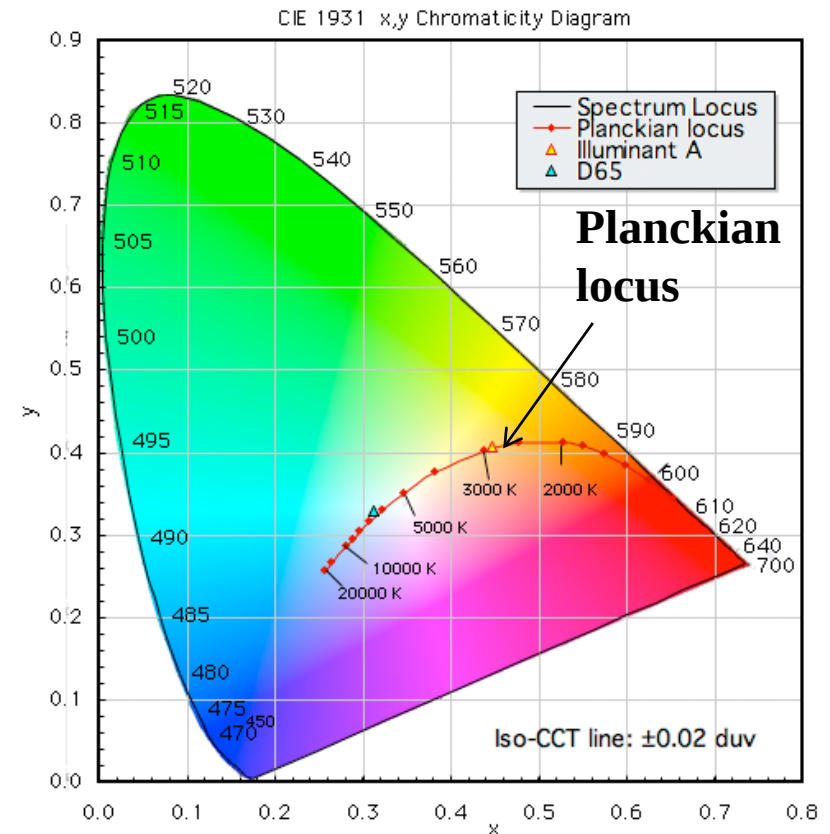
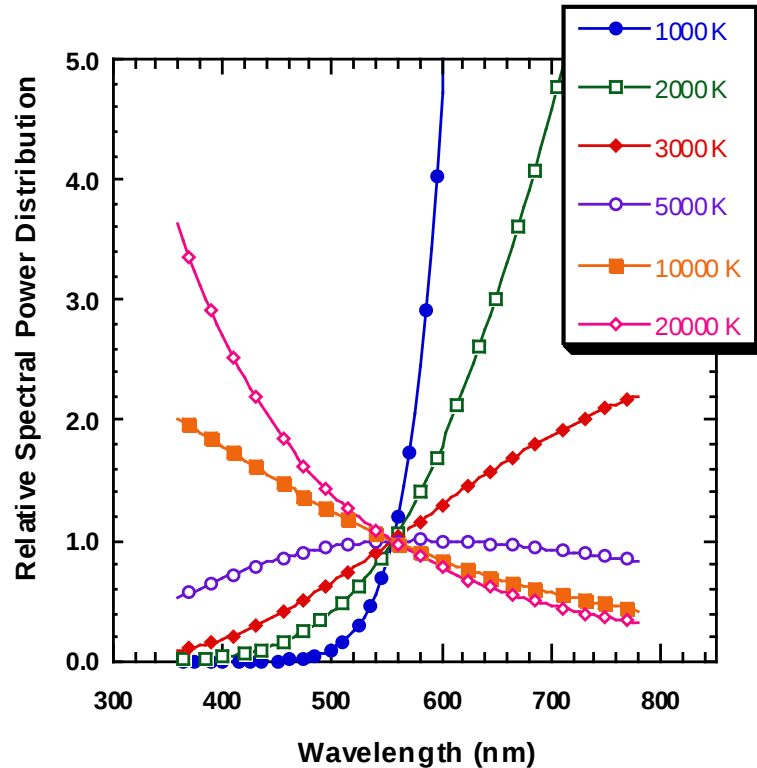


Spectral Radiance of a Planckian radiator (blackbody)

$$L_e(\lambda, T) = c_1 n^{-2} \pi^{-1} \lambda^{-5} \left[\exp\left(\frac{c_2}{n\lambda T}\right) - 1 \right]^{-1}$$



Color Temperature

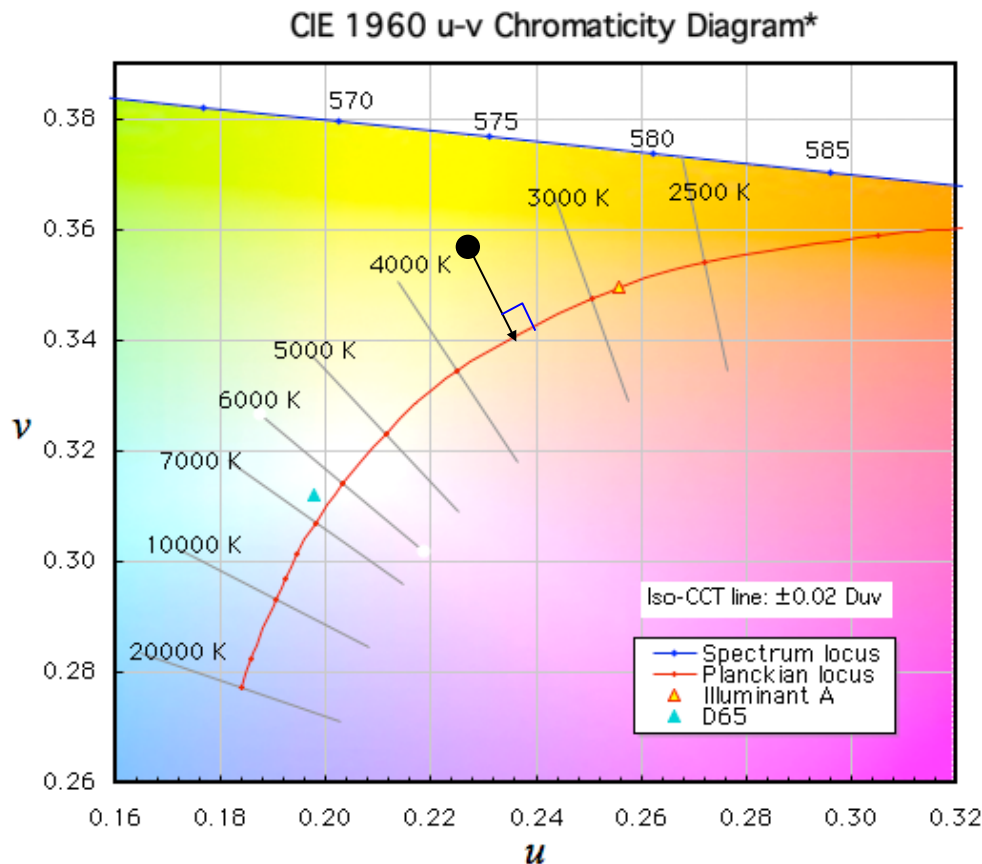


Colors along the Planckian locus are expressed by the temperature (K) of the blackbody.

Correlated Color Temperature (CCT)

Color of a light source can be expressed by the color temperature of the closest point on the Planckian locus on a uniform chromaticity diagram*.

* CIE 1960 (u, v) diagram



Symbol: T_{cp}

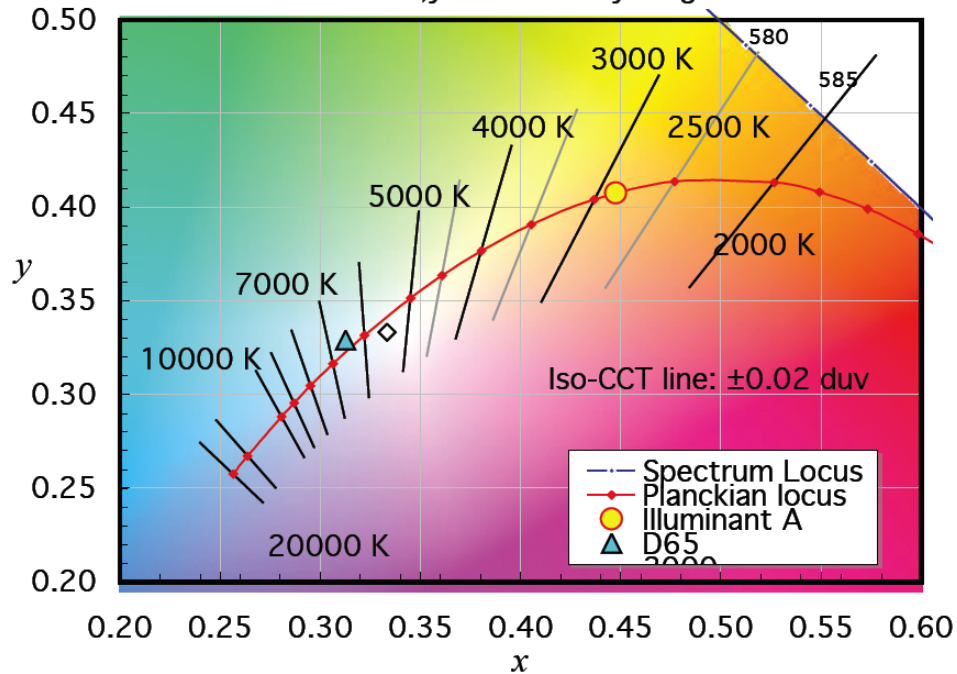
Applicable to white light sources only

CCT is based on the CIE 1960 (u, v) diagram, which is now obsolete.

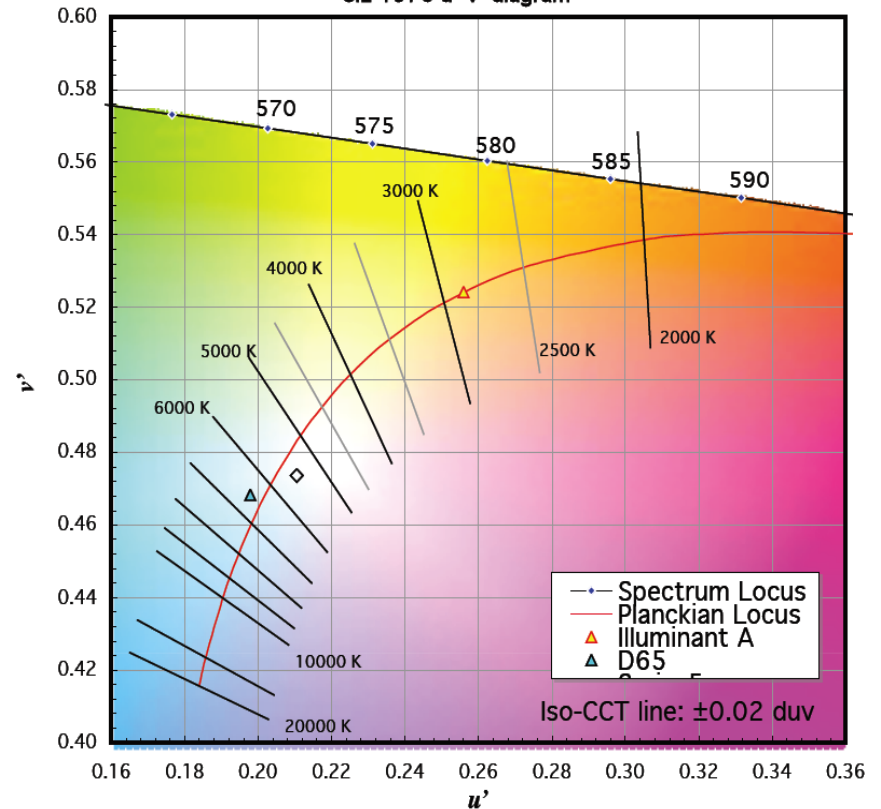
CCT is valid within distance 0.05 from the Planckian locus on the (u' , $2/3 \cdot v'$) coordinate. (CIE 15:2004)

iso-CCT lines

CIE 1931 x,y Chromaticity Diagram



CIE 1976 u'-v' diagram



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COMMISSION INTERNATIONALE DE L'ECLAIRAGE
INTERNATIONAL COMMISSION ON ILLUMINATION
INTERNATIONALE BELEUCHTUNGSKOMMISSION

CIE Central Bureau Kegelgasse 27 A-1030 Wien Austria

CIE: International Commission on Illumination

Established in 1913.

CIE defines quantities in optical metrology, publishes standards in light and lighting

CIE Central Bureau in Vienna, Austria

A. Webb (UK), President

Y. Ohno (USA), Vice President-Technical

Divisions

Division 1 Vision and Color

Division 2 Physical measurement of Light and Radiation

Division 3 Interior Environment and Lighting Design

Division 4 Lighting and Signalling for Transport

Division 5 Exterior Lighting and Other Applications

Division 6 Photobiology and Photochemistry

Division 8 Image Technology

The Meter Convention



CIPM (International Committee for Weights and Measures)

CCPR (Consultative Committee for Photometry and Radiometry)

CCPR defines units in optical metrology

Chair: T. Usuda (Japan)

WG on Key Comparisons (Y. Ohno, chair)

WG on Strategic Planning (J. Zwinkels, chair)

WG on CMC (rotates)

CCPR Key Comparisons

K1 (a,b,c) spectral irradiance

K2 (a,b,c) spectral responsivity

K3 luminous intensity

K4 luminous flux

K5 spectral diffuse reflectance

K6 spectral regular transmittance



Inter-American Metrology System (SIM)

Technical Committee of Photometry and Radiometry (TC-PR)

Chair: Maria Nadal (NIST, USA)

Key Comparisons

SIM.PR-K4: luminous flux

SIM.PR-K6: spectral regular transmittance (in preparation)

SIM.PR-S3: cryogenic radiometer, bilateral (NIST-INMETRO)

Planned

K3 luminous intensity

more

OUTLINE

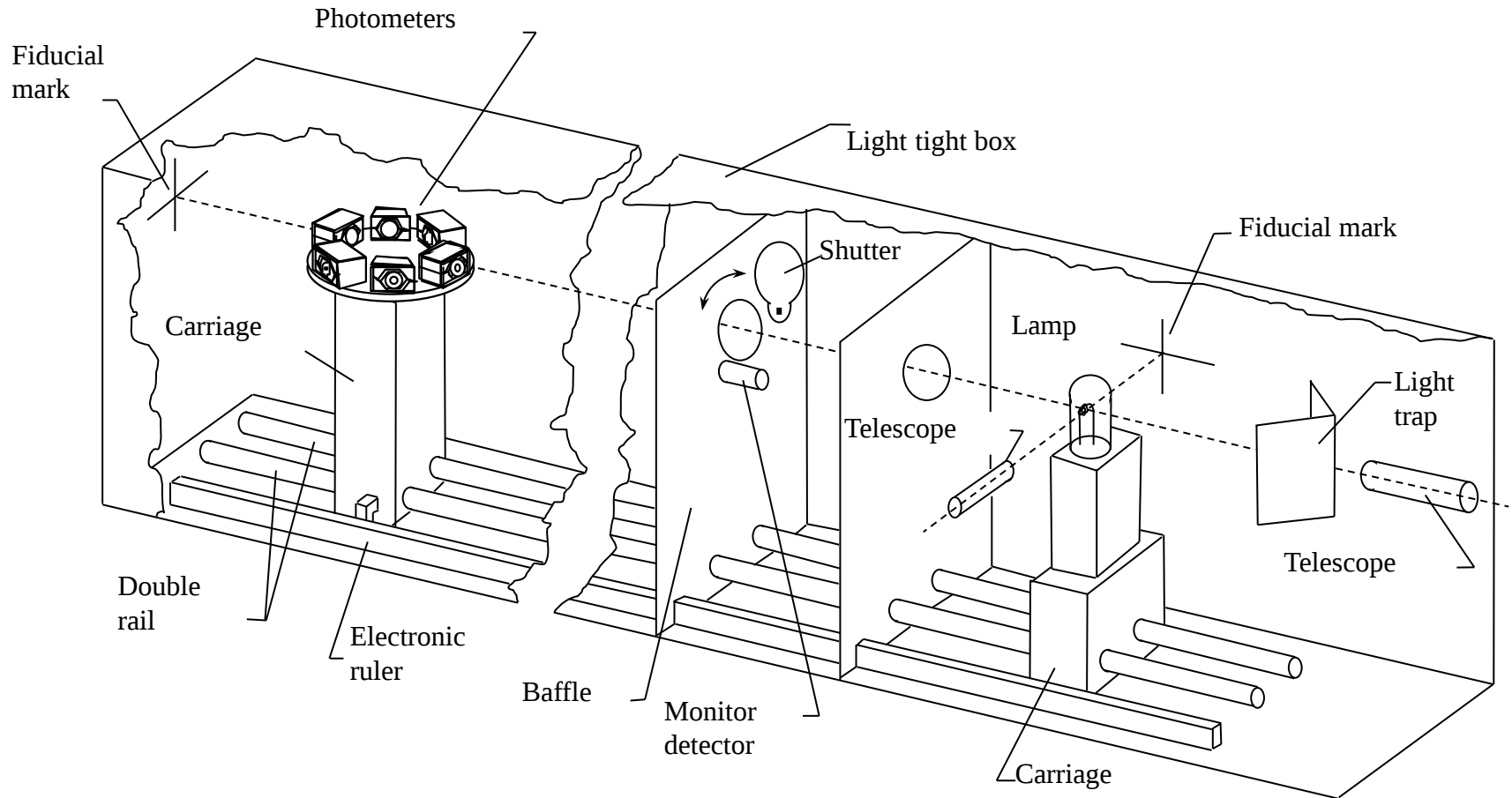
1. Optical Radiation
2. Radiometry and photometry
3. Quantities and units for optical radiation
4. Human visual response
5. Colorimetry
6. Standards organizations
- 7. Laboratory Examples**

Example 1: Measurement of Luminous Intensity



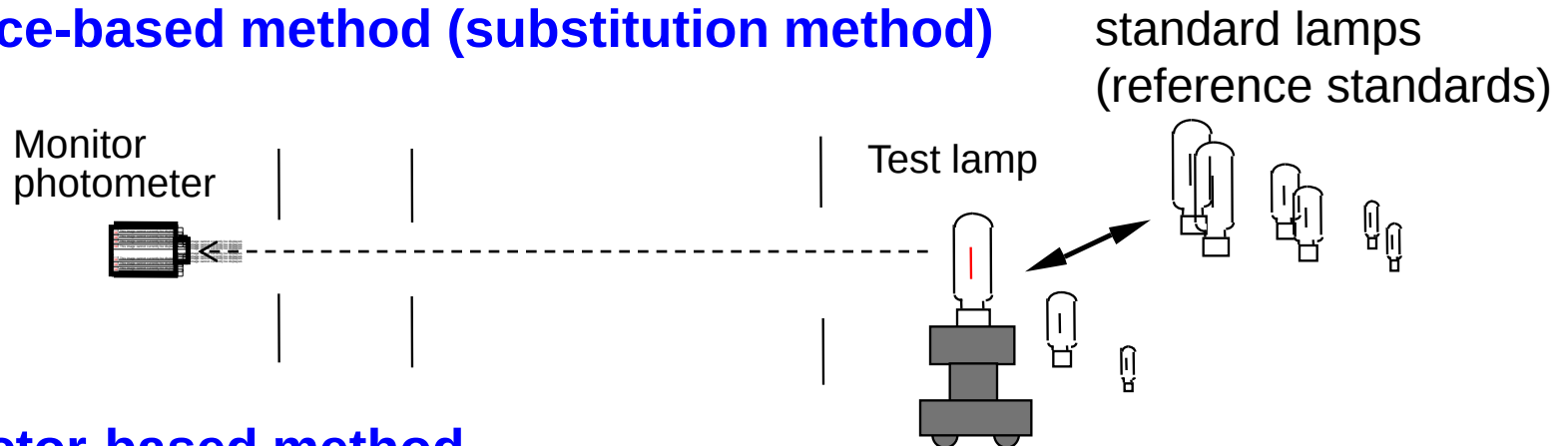
NIST photometry bench

NIST photometry bench

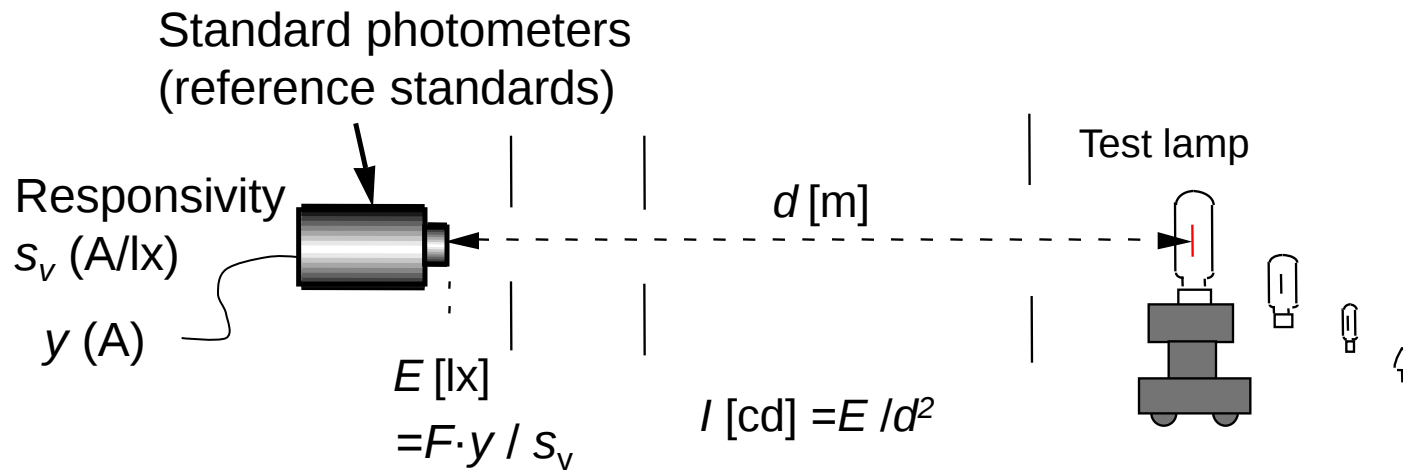


Example 1: Measurement of Luminous Intensity

Source-based method (substitution method)

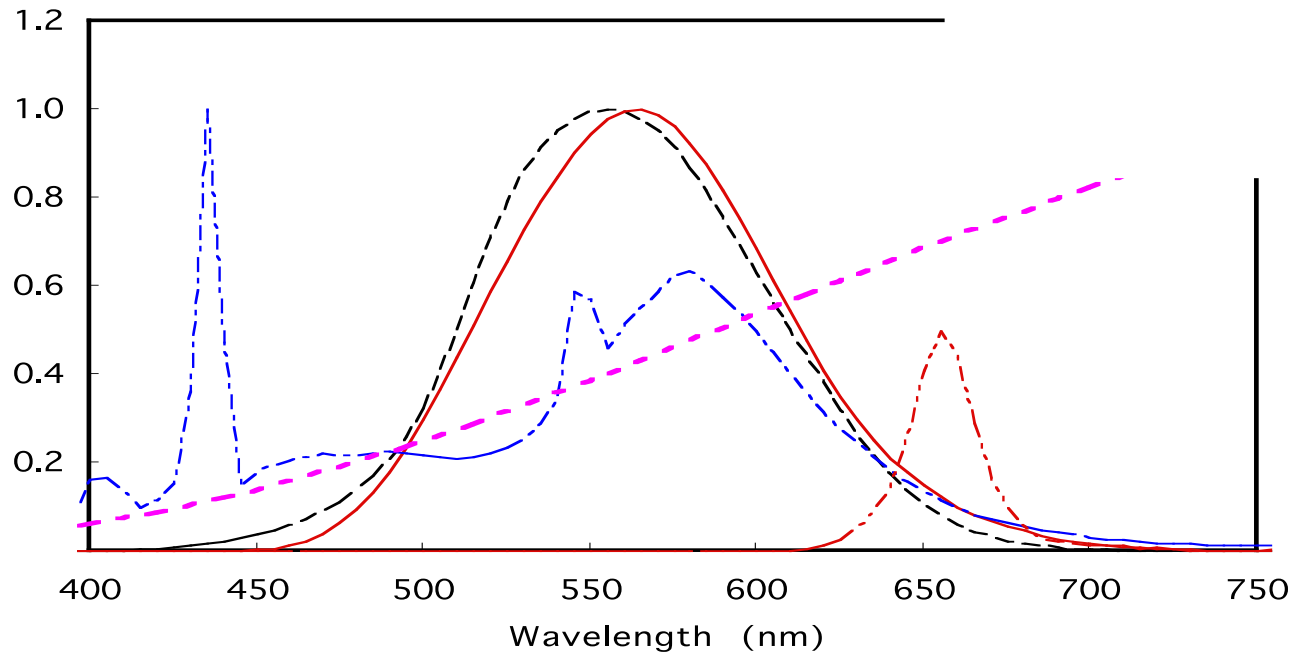


Detector-based method



F : spectral mismatch correction factor

Spectral mismatch correction



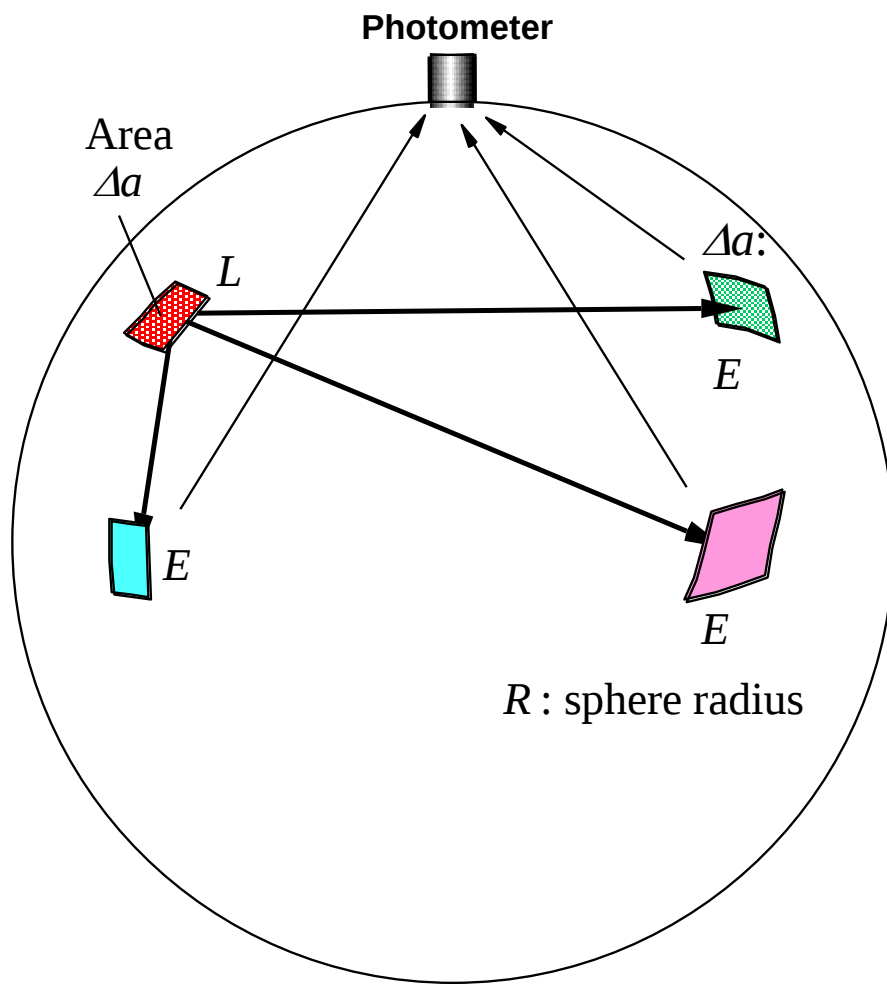
$$F(S_t, S_s) = \frac{\int_{\lambda} S_s(\lambda) R_s(\lambda) d\lambda \int_{\lambda} S_t(\lambda) V(\lambda) d\lambda}{\int_{\lambda} S_s(\lambda) V(\lambda) d\lambda \int_{\lambda} S_t(\lambda) R_s(\lambda) d\lambda}$$

$S_s(\lambda)$: SPD of standard lamp
 $S_t(\lambda)$: SPD of test lamp
 $R_s(\lambda)$: spectral responsivity of sphere system

Example 2: Measurement of Total Luminous Flux



Integrating Sphere Theory



Luminance L on a surface element creates equal illuminance ΔE all around the sphere surface

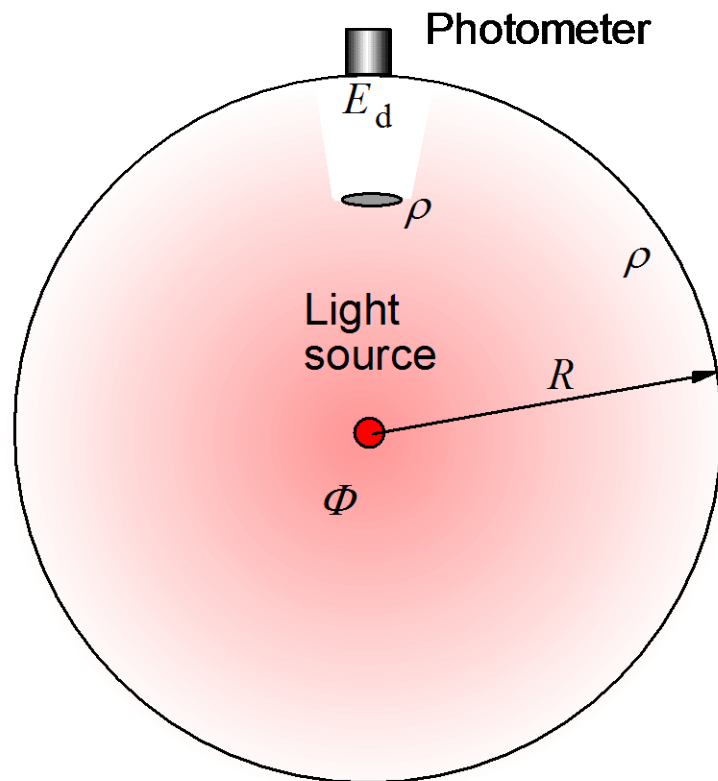
$$E = L \Delta a / 4R^2$$

The same amount of flux incident anywhere on the surface on the wall creates an equal illuminance on the photometer port.

Premise

* Sphere wall is perfectly Lambertian.

Measurement of total luminous flux with an integrating sphere



Direct illumination from the source to the detector must be shielded so that the detector measures only the flux from the sphere wall.

Flux created by interreflections:

$$\Phi (\rho + \rho^2 + \rho^3 \dots) = \Phi \frac{\rho}{1-\rho}$$

Illuminance E_d created by interreflections

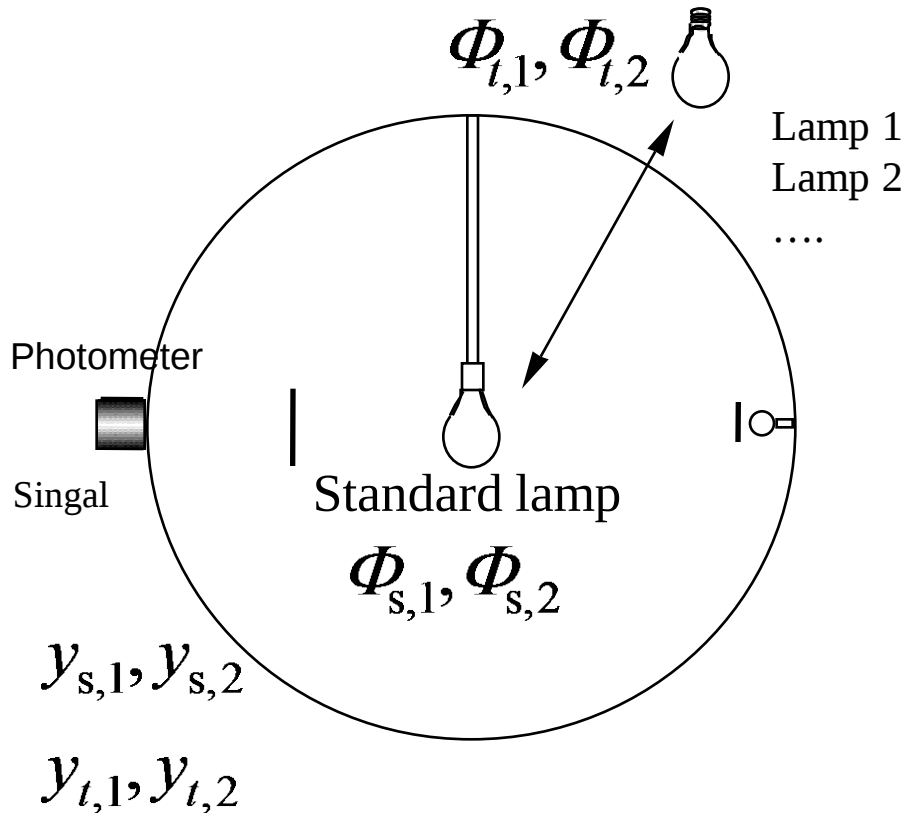
$$E_d = \frac{\Phi \rho}{1-\rho} \cdot \frac{1}{4 \pi R^2}$$

This equation indicates:

1. The sphere throughput (E_d / Φ) is strongly dependent on ρ .
(E_d at $\rho=0.98$ is 10 times larger than at $\rho=0.8$)
2. The sphere throughput will be very sensitive to ρ when ρ is close to 1.
(0.1 % change of ρ at $\rho=0.98$ causes 5 % change of the signal.)

Substitution Method

Test lamp



The sphere responsivity is calibrated as

$$R_s = \text{ave} \left(\frac{y_{s,i}}{\Phi_{s,i}} \right) \quad [\text{A/lm}]$$

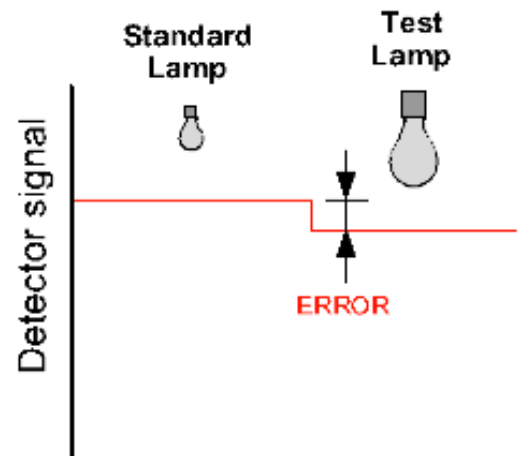
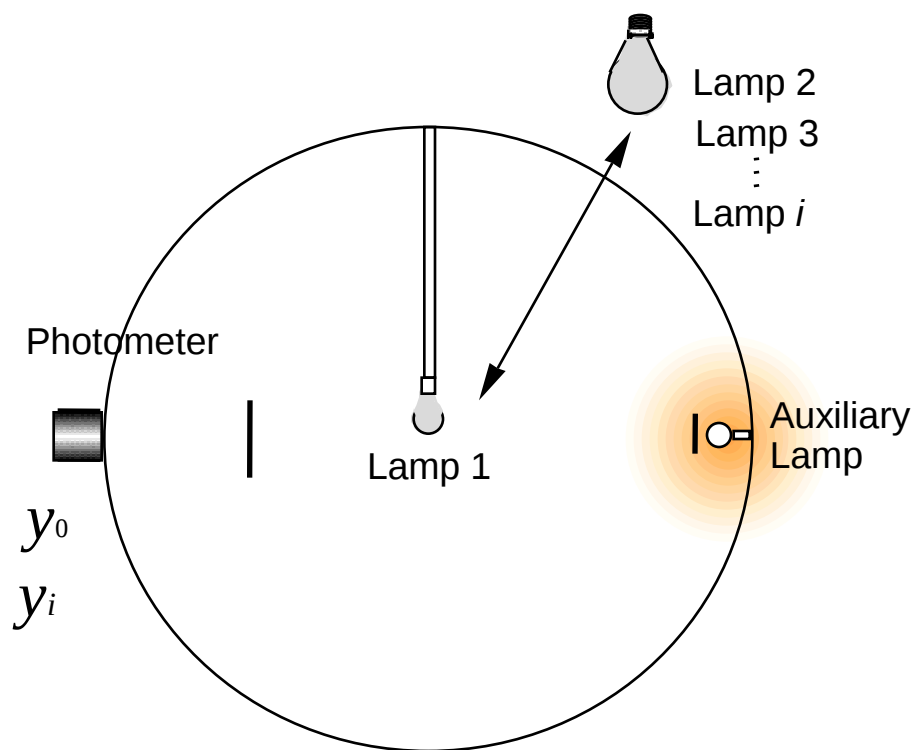
Luminous flux of test lamp $\Phi_{t,i}$ is obtained by

$$\Phi_{t,i} = \frac{y_{t,i}}{R_s} \cdot \frac{F}{\alpha} \quad [\text{lm}]$$

α : self-absorption factor

F : spectral mismatch correction factor

Self-absorption Correction



With the auxiliary lamp turned on, measure the detector signals :

y_0 : Signal with standard lamp

y_i : Signal with test Lamp

Self-absorption Factor α_i for Lamp i is obtained by

$$\alpha_i = \frac{y_i}{y_0}$$

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- Y. Ohno, Practical Use and Calculation of CCT and Duv, LEUKOS, 10:1, 47-55, DOI: 10.1080/15502724.2014.839020 (2013).

Questions

Are the following statements correct or not (Yes/No)?

1. The visible region is 360 nm to 830 nm (Yes/No).
2. Radiant energy in radiometry corresponds to luminous flux in photometry. (Yes/No)
3. Luminance of a surface is the same at any distance it is measured. (Yes/No)
4. Integrating sphere is used for luminous intensity measurement. (Yes/No)
5. Incandescent lamps have a correlated color temperatures of about 5000 K. (Yes/No)